

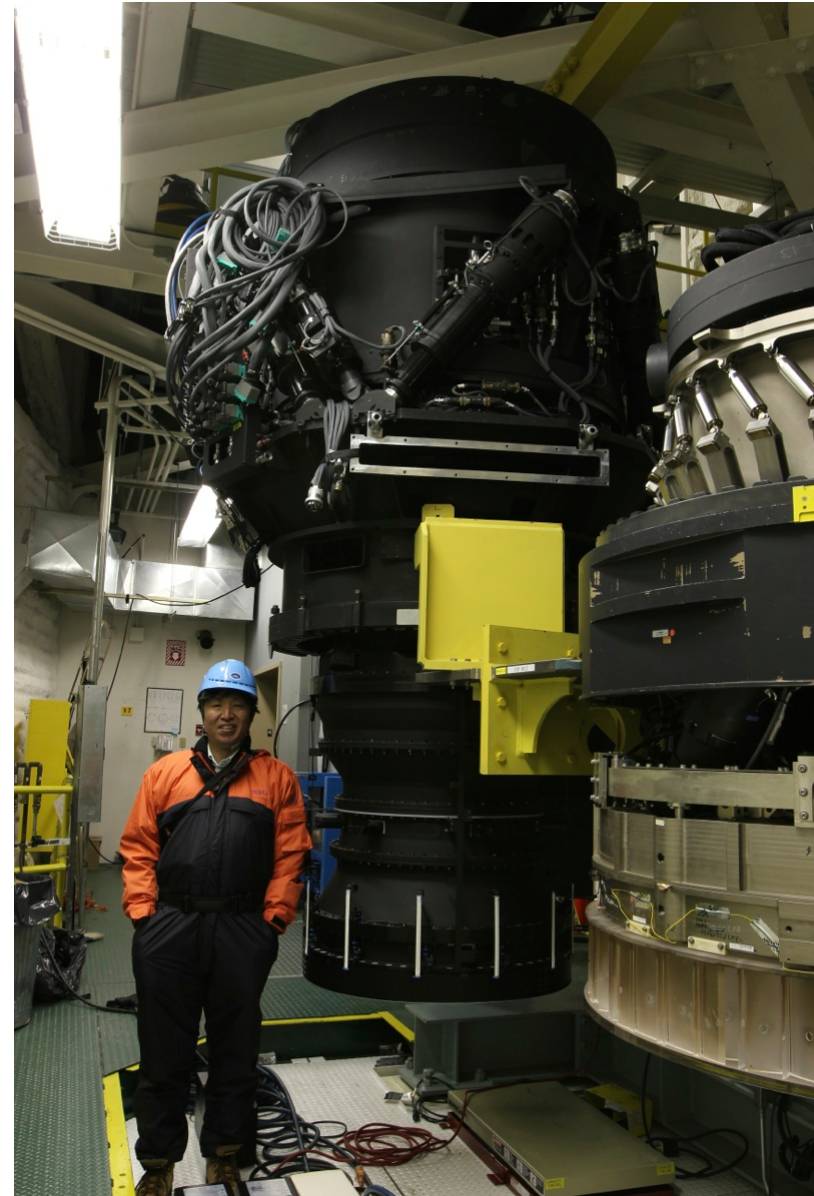
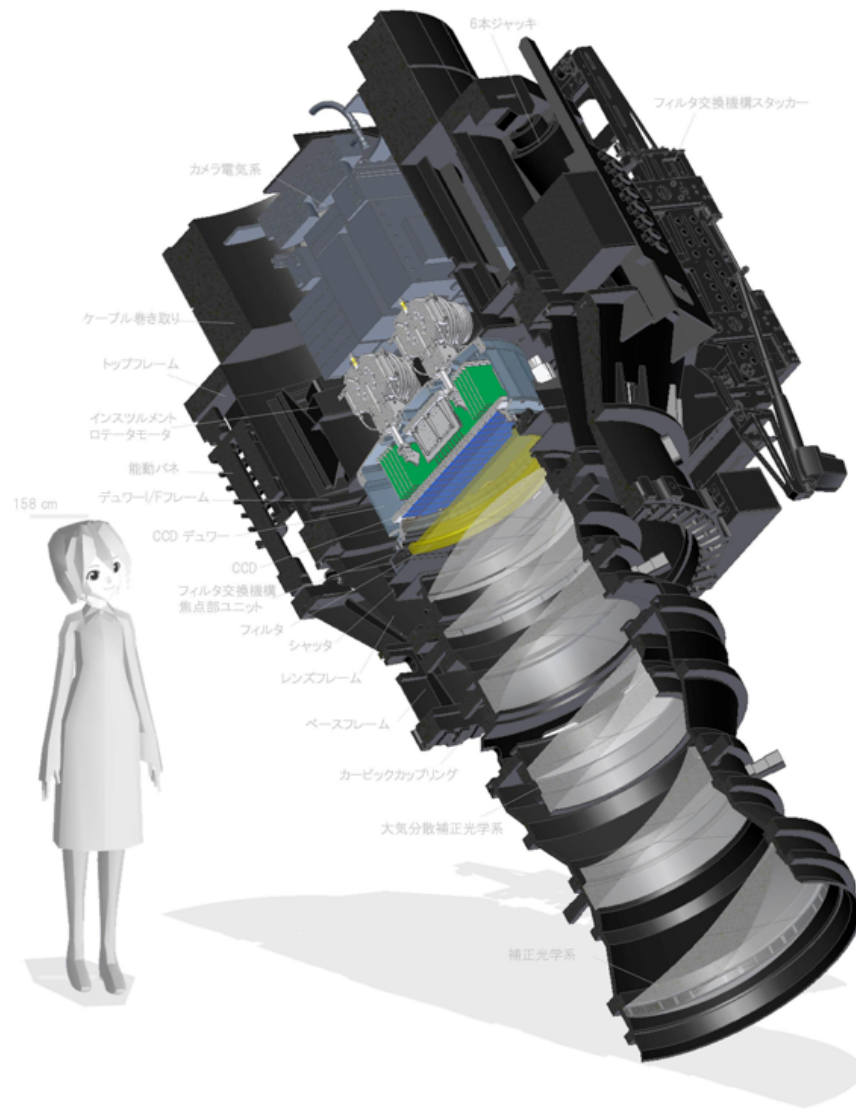
Hyper Suprime-Cam

現状報告





Hyper Suprime-Cam





Engineering First Light

- 2012/08/28 ~
 - Auto Guider, Pointing Analysis System, Mirror Analysis Systemのsoftware/hardwareの機能確認
- ほぼOK。次は性能の確認
- Filter Exchange Unitは所定の試験回数をこなす時間が足りず、安全をみて搭載を見送った



Hyper Suprime-Cam

今後の試験観測の予定

1/24 - 2/4

3/11 - 3/21

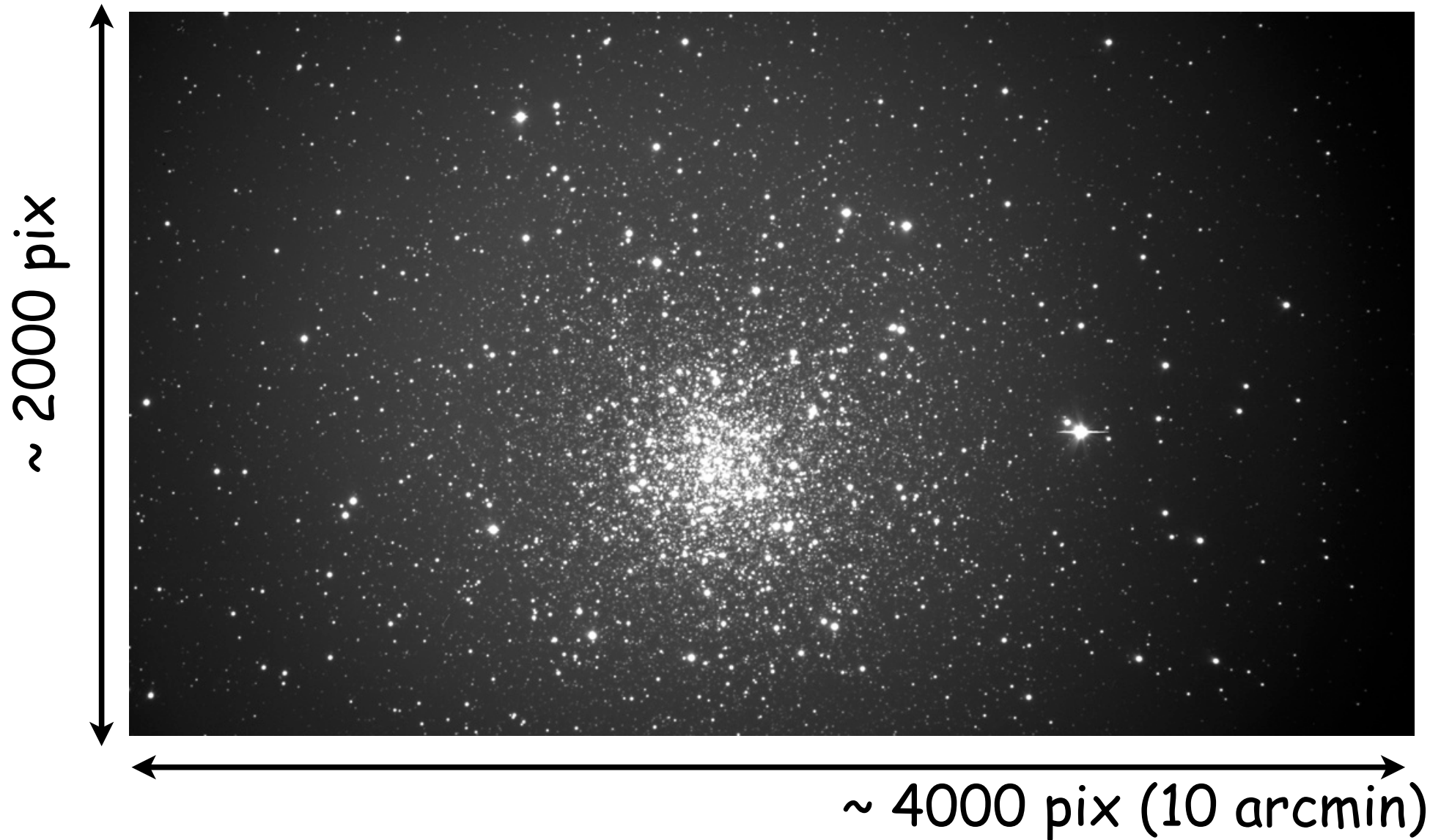
5/16 - 5/20

6/6 - 6/13



Main Camera Image

M56 10 sec r-band $\sim 0''.58$ FWHM

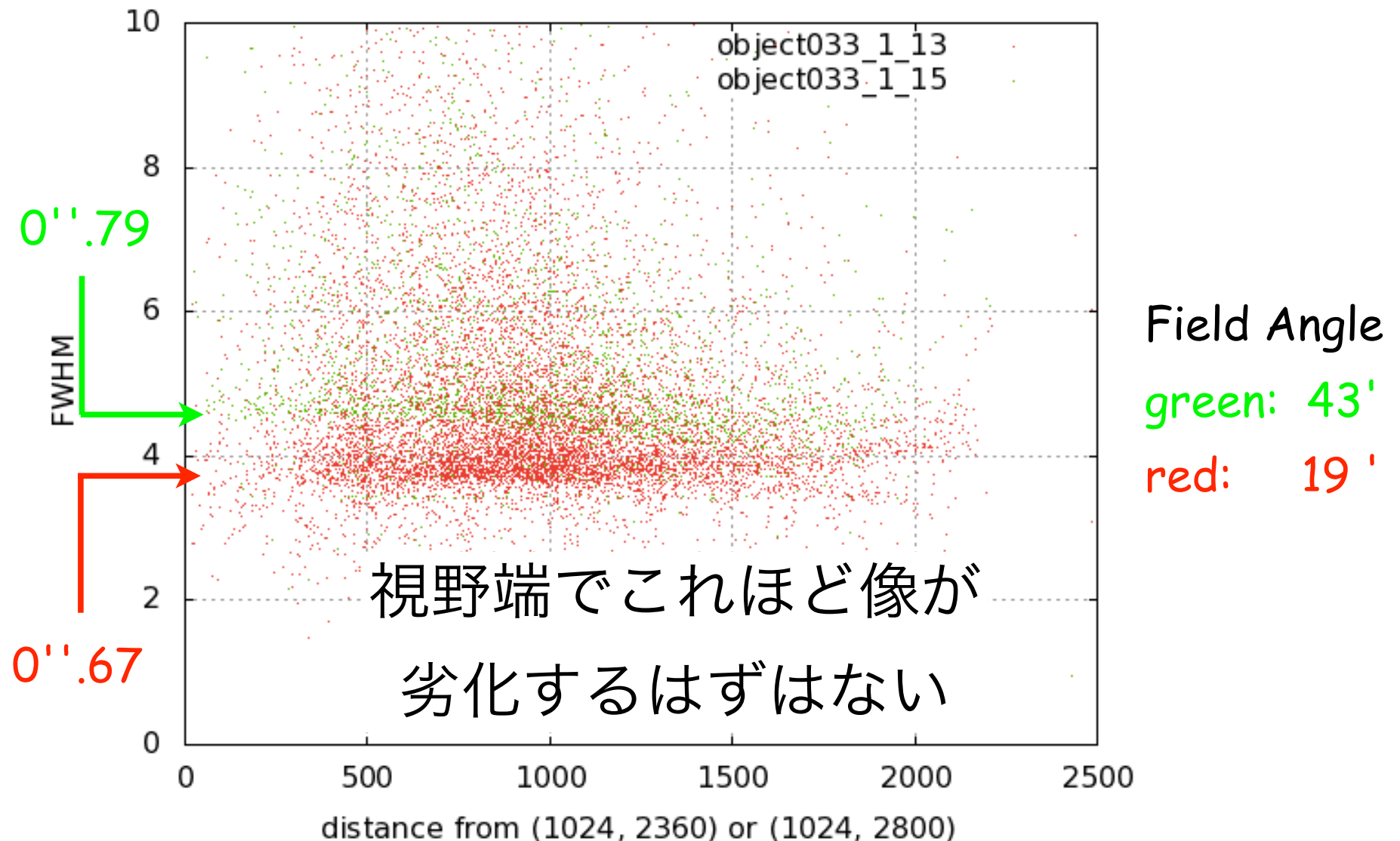




Jan2013 Engineering Plan

24	PA, Open Track, <u>ADC</u> link	M
25-26	SH Coordinate Check	M
27-28	PM Control using SH	M
29-30	HSC SC AG Function Test	MK
1/31	Spare	K
2/1-4	(WFC+Camera) Tilt& misc Measurement	K

Main Camera Image Analysis





戦略枠プロジェクト

Wide-field imaging with Hyper Suprime-Cam: Cosmology and Galaxy Evolution *A Strategic Survey Proposal for the Subaru Telescope*

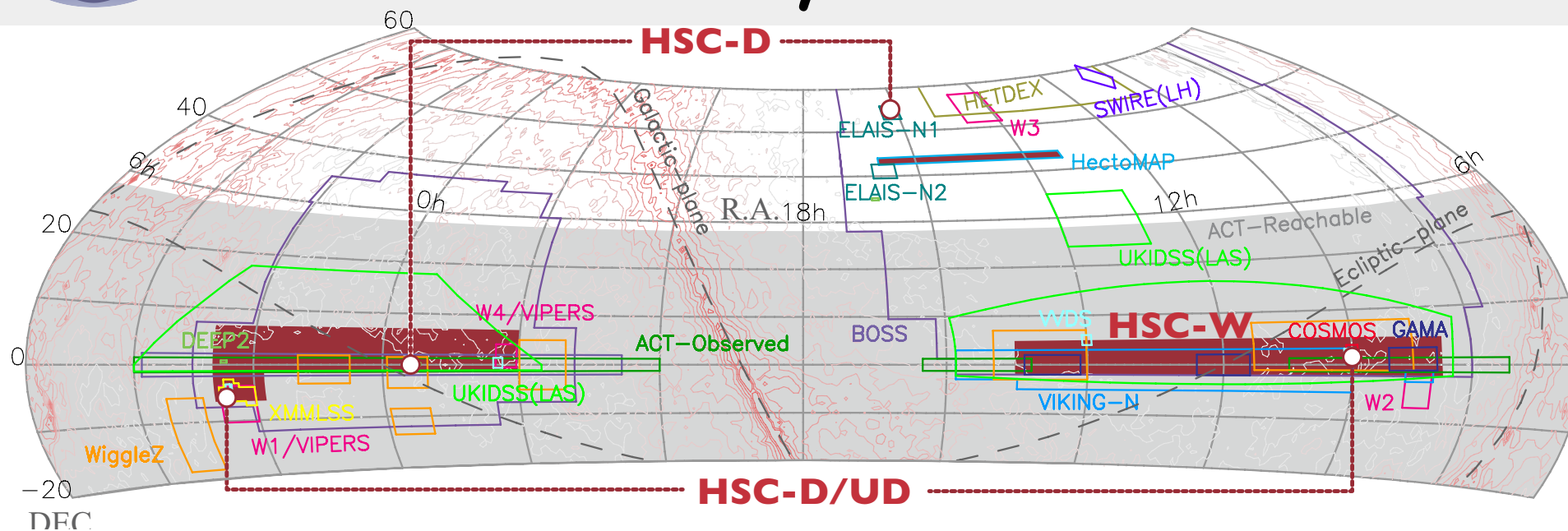
PI: Satoshi Miyazaki (NAOJ)

Co-PI: Ikuru Iwata (NAOJ)

The HSC collaboration team¹: S. Abe⁽¹⁾, H. Aihara^{*(2),(3)}, M. Akiyama⁽⁴⁾, K. Aoki⁽⁵⁾, N. Arimoto^{*(5)}, N. A. Bahcall⁽⁶⁾, S. J. Bickerton⁽³⁾, J. Bosch⁽⁶⁾, K. Bundy^{†(3)}, C. W. Chen⁽⁷⁾, M. Chiba^{†(4)}, T. Chiba⁽⁸⁾, N. E. Chisari⁽⁶⁾, J. Coupon⁽⁷⁾, M. Doi⁽²⁾, M. Enoki⁽⁹⁾, S. Foucaud⁽¹⁰⁾, M. Fukugita⁽³⁾, H. Furusawa^{†(5)}, T. Futamase⁽⁴⁾, R. Goto⁽²⁾, T. Goto⁽¹¹⁾, J. E. Greene⁽⁶⁾, J. E. Gunn^{†(6)}, T. Hamana^{†(5)}, T. Hashimoto⁽²⁾, M. Hayashi⁽⁵⁾, Y. Higuchi^{(2),(5)}, C. Hikage⁽¹²⁾, J. C. Hill⁽⁶⁾, P. T. P. Ho^{*(7)}, B. C. Hsieh⁽⁷⁾, K. Y. Huang^{†(7)}, H. Ikeda⁽¹³⁾, M. Imanishi⁽⁵⁾, N. Inada⁽¹⁴⁾, A. K. Inoue⁽¹⁵⁾, W.-H. Ip⁽¹⁾, T. Ito⁽⁵⁾, K. Iwasawa⁽¹⁶⁾, M. Iye⁽⁵⁾, H. Y. Jian⁽¹⁷⁾, Y. Kakazu⁽¹⁸⁾, H. Karoji⁽³⁾, N. Kashikawa⁽⁵⁾, N. Katayama⁽³⁾, T. Kawaguchi⁽¹⁹⁾, S. Kawanomoto⁽⁵⁾, I. Kayo⁽²⁰⁾, T. Kitayama⁽²⁰⁾, G. R. Knapp⁽⁶⁾, T. Kodama⁽⁵⁾, K. Kohno⁽²⁾, M. Koike⁽⁵⁾, E. Kokubo⁽⁵⁾, M. Kokubo⁽²⁾, Y. Komiyama⁽⁵⁾, A. Konno⁽²⁾, Y. Koyama⁽⁵⁾, C. N. Lackner⁽³⁾, D. Lang⁽⁶⁾, A. Leauthaud^{†(3)}, M. J. Lehner⁽⁷⁾, K.-Y. Lin⁽⁷⁾, L. Lin⁽⁷⁾, Y.-T. Lin^{†(7)}, C. P. Loomis⁽⁶⁾, R. H. Lupton^{†(6)}, P. S. Lykawka⁽²¹⁾, K. Maeda⁽³⁾, R. Mandelbaum^{†(22)}, Y. Matsuda⁽⁵⁾, K. Matsuoka^{(13),(23)}, Y. Matsuoka⁽¹²⁾, S. Mineo⁽²⁾, T. Minezaki⁽²⁾, H. Miyatake⁽⁶⁾, R. Momose⁽²⁾, A. More⁽³⁾, S. More⁽³⁾, T. J. Moriya⁽³⁾, T. Morokuma^{†(2)}, H. Murayama^{*(3)}, K. Nagamine⁽²⁴⁾, T. Nagao^{†(23)}, S. Nagataki⁽²³⁾, Y. Naito⁽²⁾, K. Nakajima⁽²⁾, F. Nakata⁽⁵⁾, H. Nakaya⁽⁵⁾, T. Namikawa⁽²⁾, C.-C. Ngeow⁽¹⁾, T. Nishimichi⁽³⁾, H. Nishioka⁽⁷⁾, A. J. Nishizawa^{†(3)}, K. Nomoto⁽³⁾, M. Oguri^{†(3)}, A. Oka⁽²⁾, N. Okabe⁽⁷⁾, S. Okamoto⁽²⁵⁾, S. Okamura⁽²⁶⁾, J. Okumura⁽²³⁾, S. Okumura⁽²⁷⁾, Y. Okura⁽⁵⁾, Y. Ono⁽²⁾, M. Onodera⁽²⁸⁾, K. Ota⁽²³⁾, M. Ouchi^{†(2)}, S. Oyabu⁽¹²⁾, P. A. Price⁽⁶⁾, R. Quimby⁽³⁾, C. E. Rusu^{(2),(5)}, S. Saito⁽²⁹⁾, T. Saito⁽³⁾, Y. Saitou⁽³⁰⁾, M. Sato⁽¹²⁾, T. Shibuya⁽⁵⁾, K. Shimasaku^{†(2)}, A. Shimono⁽³⁾, S. Shinogi⁽²⁾, M. Shirasaki⁽²⁾, J. D. Silverman⁽³⁾, D. N. Spergel^{*(6),(3)}, M. A. Strauss^{†(6)}, H. Sugai⁽³⁾, N. Sugiyama^{(12),(3)}, D. Suto⁽²⁾, Y. Suto^{*(2)}, K. Tadaki⁽²⁾, M. Takada^{†(3)}, R. Takahashi⁽³¹⁾, S. Takahashi⁽⁵⁾, T. Takata⁽⁵⁾, T. T. Takeuchi⁽¹²⁾, N. Tamura⁽³⁾, M. Tanaka⁽⁵⁾, M. Tanaka^{†(3)}, M. Tanaka⁽⁴⁾, Y. Taniguchi⁽¹³⁾, A. Taruya⁽²⁾, T. Terai⁽⁵⁾, Y. Terashima⁽¹³⁾, N. Tominaga⁽³²⁾, J. Toshikawa⁽³⁰⁾, T. Totani⁽²³⁾, M. Tsai⁽¹⁾, E. L. Turner^{*(6)}, Y. Ueda⁽²³⁾, K. Umetsu⁽⁷⁾, Y. Urata^{†(1)}, Y. Utsumi⁽⁵⁾, B. Vulcani⁽³⁾, K. Wada⁽³³⁾, S.-Y. Wang⁽⁷⁾, W.-H. Wang⁽⁷⁾, T. Yamada⁽⁴⁾, Y. Yamada⁽⁵⁾, K. Yamamoto⁽³⁴⁾, H. Yamanoi⁽⁵⁾, C.-H. Yan⁽⁷⁾, N. Yasuda^{†(3)}, A. Yonehara⁽³⁵⁾, F. Yoshida^{†(5)}, N. Yoshida⁽²⁾, M. Yoshikawa⁽³⁶⁾, S. Yuma⁽²⁾ (1) NCU, Taiwan (2) Tokyo (3) Kavli IPMU (4) Tohoku (5) NAOJ (6) Princeton (7) ASIAA (8) Nihon (9) Tokyo Keizai (10) NTNU, Taiwan (11) DARK, Copenhagen (12) Nagoya (13) Ehime (14) NNCT (15) Osaka Sangyo (16) Barcelona (17) NTU, Taiwan (18) Chicago (19) Tsukuba (20) Toho (21) Kinki (22) CMU (23) Kyoto (24) Las Vegas (25) KIAA, China (26) Hosei (27) JSGA (28) ETH (29) Berkeley (30) GUAS (31) Hirosaki (32) Konan (33) Kagoshima (34) Hiroshima (35) Kyoto Sangyo (36) JAXA

- WLを使いDEを前人未踏の精度で計測
- 日本がリードしてきた初期銀河宇宙研究を発展

Survey Field



Wide 1500 deg²

Deep 28 deg²

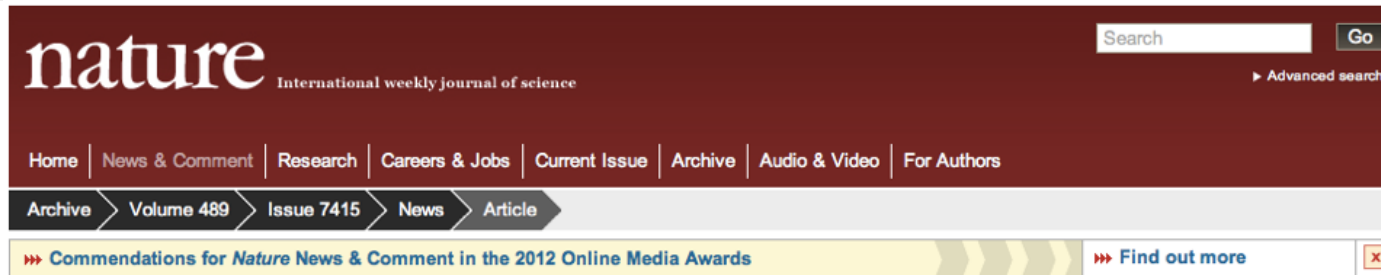
UDeep 4 deg²

10 times wider, 1 mag
deeper, 20 % sharper
than CFHLS-Wide

大規模サーベイ、解析遂行にはcommunity effortの
結集が必要。Wのみ、D/UDのみでは実行不可能。



Nature も注目!



NATURE | NEWS

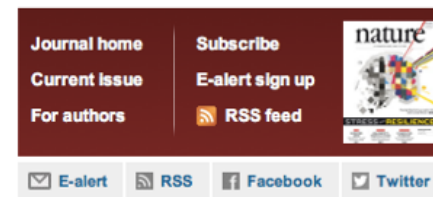
Cameras to focus on dark energy

A pair of detectors that measure minute distortions in images of distant galaxies will probe the riddle of cosmic acceleration.

Eric Hand

12 September 2012

Even the best pictures of a distant galaxy are a bit lopsided. But this is an attribute, not a bug. Because mass distorts space-time, light coming from distant galaxies is bent as it passes through intervening shoals of invisible matter, leaving the images of these distant objects minutely sheared and stretched.



Nature 489, 190-191
(13 September 2012)

"One of the new efforts will use the Hyper Suprime-Cam (HSC), which achieved first light on 28 August on the 8.2-metre Subaru telescope in Hawaii. By 2018, it aims to have imaged 10 million galaxies over a 1,500-square-degree-swathe of the sky ... A rival project DES ..."

Subaru UM 2013/01/15

戦略枠プロポーザル

第1段階審査 合格

第2段階審査

開かれた(研究)組織作り

ハワイ観測所における体制作り

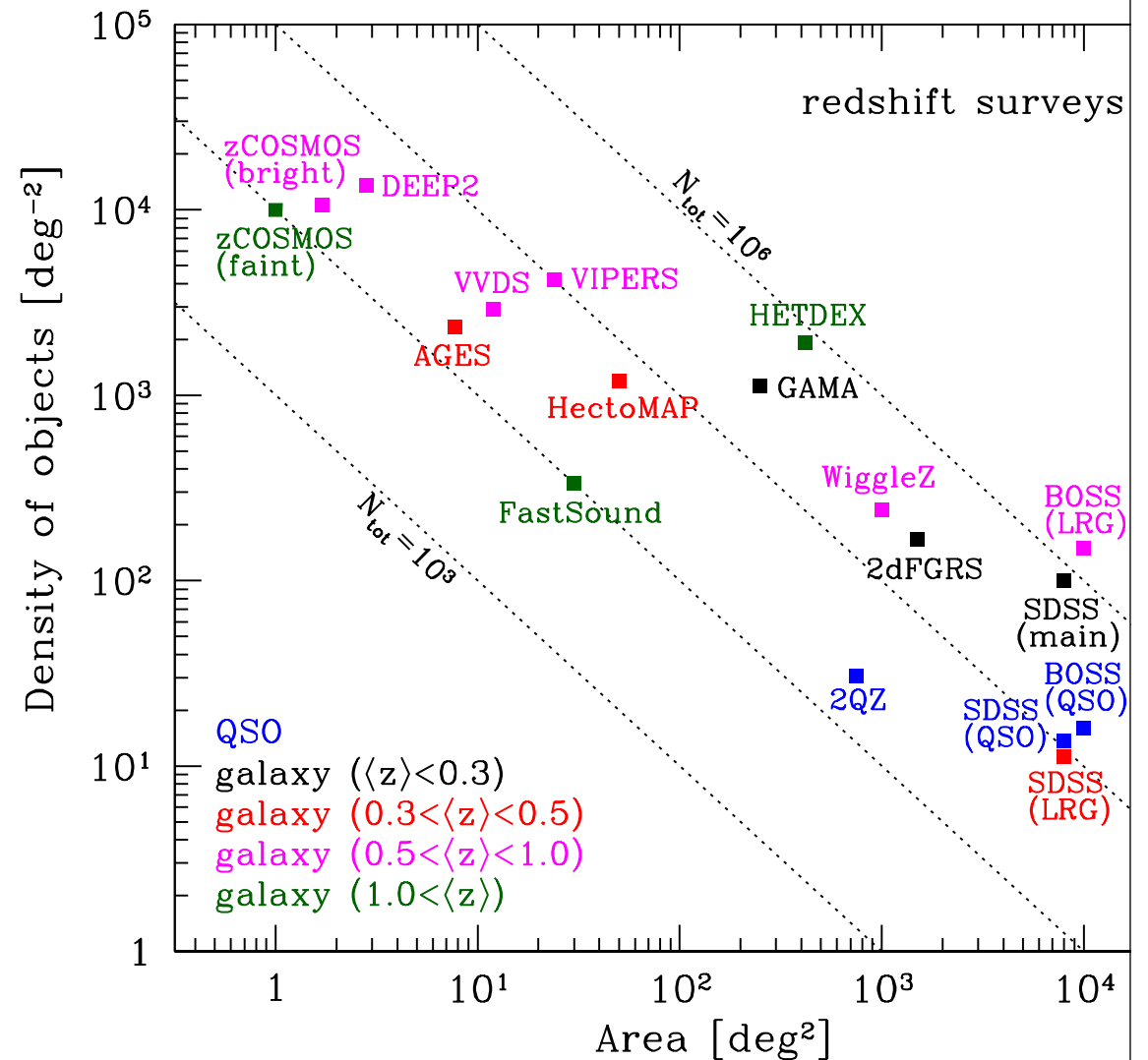
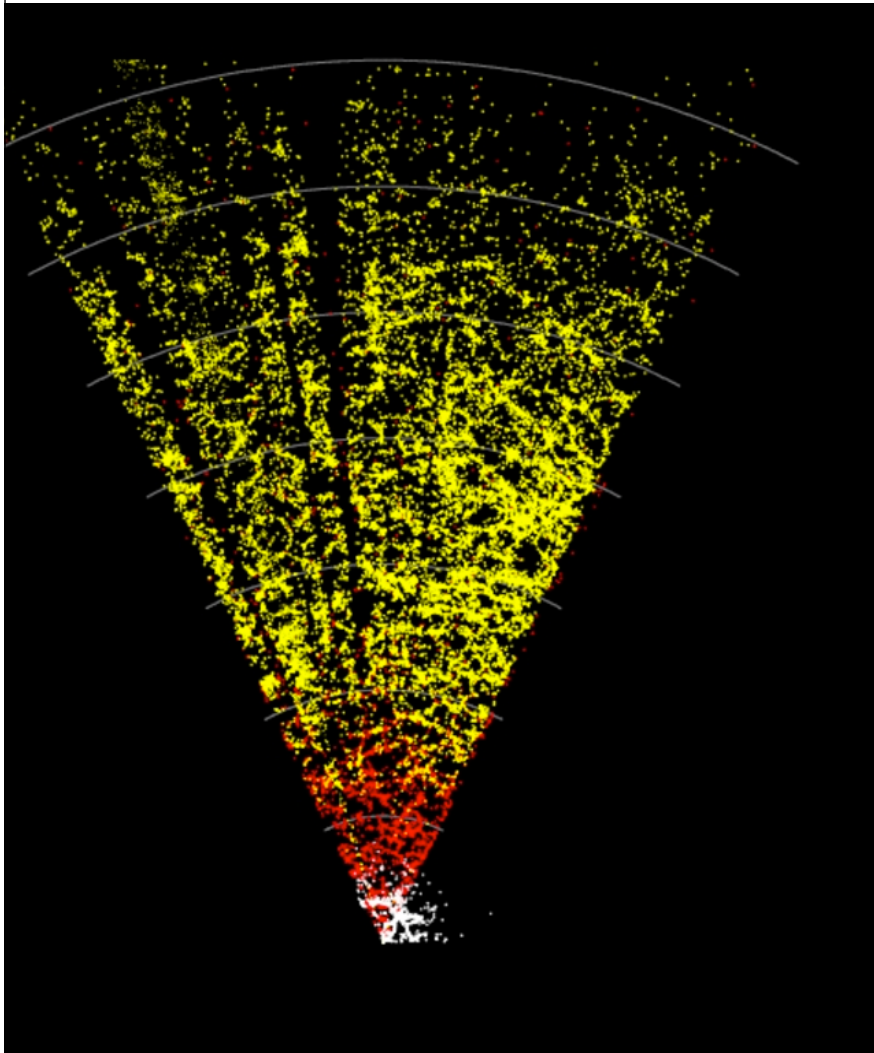
第3段階審査 3月にヒアリング

第4段階審査 最終提案書

他プロジェクトとの協力

- Spitzer COSMOS survey (Peter Capak)
- Ultra Deep COSMOSデータ(BBのみ?)と
Spitzerデータとの交換
- HectoMAP (Margaret Geller)
 - Wide (50平方度のみ)のDark Matter Halo
Catalogと分光データの交換($r < 21.3$ dense)
- U-band
 - Palomar ? CFHT ? DECam ? 本当に必要 ?

HectoMAP



白 CfA 赤 Sloan 黄 HectoMAP



University of Hawaii ?

- Meeting with JT & RJW on 2012/12
- SCの時のように組織だったsurvey計画はまだない
- 様子見だが、データが出始めると本気で考え始めるだろう
- DirectorがCOSMOS FieldのデータをHSCで取らないかと言ってきている
- UH時間でより深く撮ってもらい、shareする？
- 2013/10までに協力関係を整理したい



HSC Data Release Plan (draft)

- Raw data automatically release for public in 18 mo.
- Reduced image and Catalog release within the HSC collaboration
 - 1 st release 18 mo. from the beginning of the survey: every 12 mo. from the 2 nd
 - 18 mo. after the survey completion open globally
- Part of the catalog items release globally from the beginning ? (Position & BB mag ...)

Data Amount

- Basic Photometric Catalog
 - Single Filter: 474 Byte, Combined 1820 Byte
 - $(474 \times 5 + 1820) \sim 4 \text{ kByte/object}$
 - $\sim 50 \times 60 \times 60 \times 1400 \sim 252 \text{ M object}$
 - $\sim 1 \text{ TByte total}$
- + Weak Lensing Catalog a few TByte
- Stacked Image
 - 6×10^{11} pixels (1400 deg^2)
 - $6 \times 10^{11} \times 4 \text{ byte} \times 5 \text{ color} \sim 10 \text{ TByte}$

Image will be divided into sky patch

Distribution media: Hard Drive

- Local Database Browser
- Software Libraries to allow a combined access to catalog and stacked image for further independent analysis
- Users who want to manipulate raw images are encouraged to work closely with the HSC project's data analysis team

- Data and the software distributed through SMOKA channel
- Development of the data browsing and retrieval tools are shared by SMOKA group
- In early phase of the survey, on line distribution might be enough.



Open Use Support

- Contributing to the observatory's effort
- Provide stacked images and crude catalog made on the fly (during the observing run)
- Release data analysis software package (subset of HSC pipeline) and provide the user supports

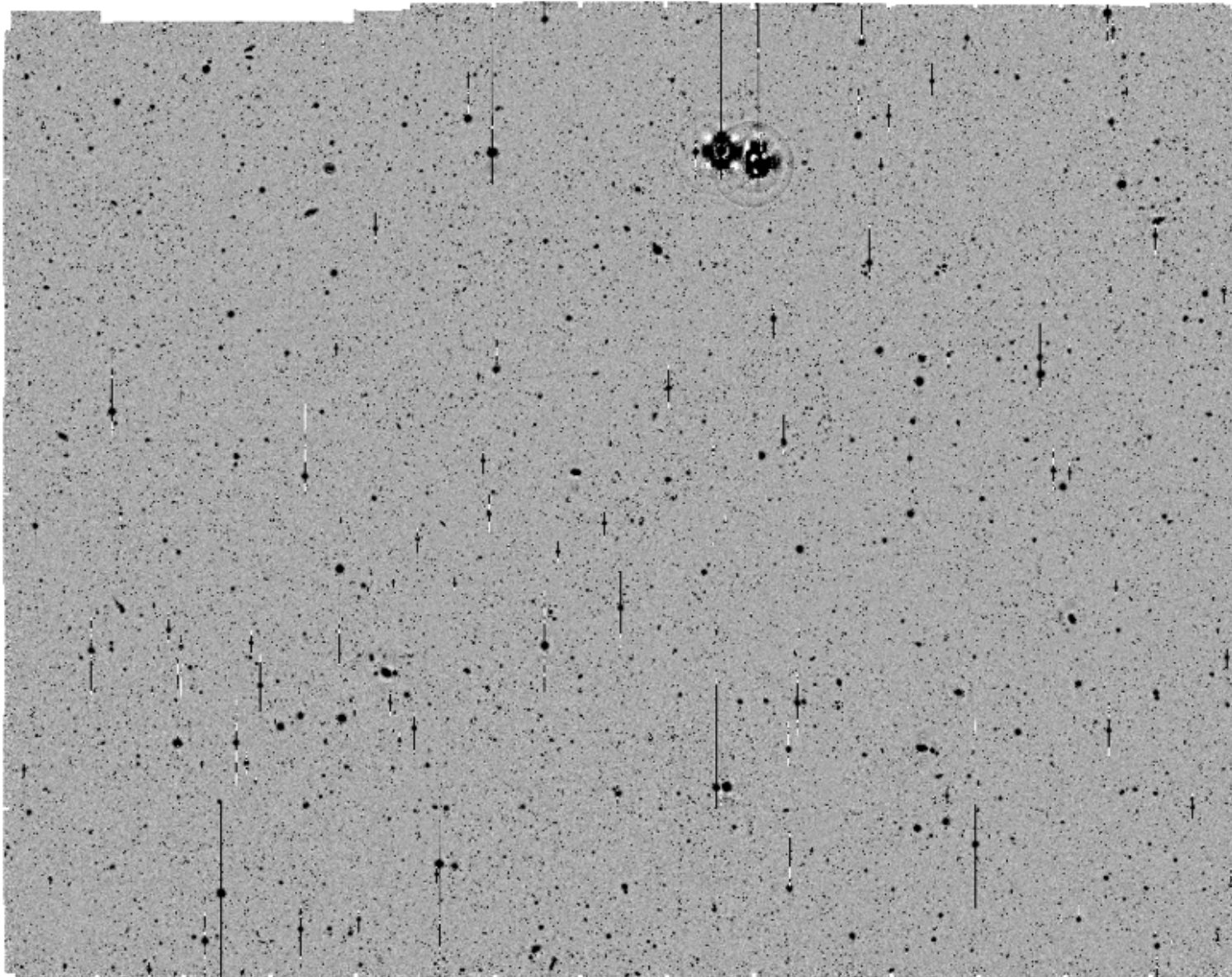
Current estimates of times to generate a mosaic-stacked image:

- Single \$3,500 PC (Linux)
- 0.5 h exposure (Rc 4 visits)

~ 2 hour

Real Time X 4₁₉

Mosaic-Stacked Image



- ✓ • 2012/08 First Light
- ~ 2013/07 Commissioning Run
- (2013/08~ Legacy Survey (5 yr) and Open Use)